## Breakout rooms

### Solution Now go to one of the following breakout rooms

You can swap later. Material from all rooms (1-5) will be provided later.

I. Relic density (standard)

- 2. Relic density (dark sector, coupled Boltzmann eqs.)
- 3. Gamma-ray (and other CR) spectra
- 4. J-factors
- 5. Neutrino signals
- 6. General technical support [installation/coding/shell usage/...]
- You can start right away with problem I and II above...

See also the link to the tutorial from the ISAPP homepage

#### Actively help each other while Joakim and me go around !

## Dark Sector relic density

Let's have a look at one of the example programs, demonstrating relic density calculations in secluded dark sectors

#### **Problem III:**

 $\label{eq:copy} Copy \ examples / aux / oh2 \_ dark\_sector.f \ \ to \ your \ private \ directory$ 

- 1. Make sure that you can compile and run this main program
- 2. Try to understand what the program does, and how
- 3. Produce a plot corresponding to one of the curves in Fig. 1 of 2007.03696.
- 4. Replace dsrdxi to explore the effect of a constant dark sector to visible sector temperature ratio
- 5. Replace dsgivemodel\_vdSIDM... to explore the effect of changing the dark sector particle content (masses and d.o.f.)

# Coupled Boltzmann equations

Let's have a look at one of the example programs, demonstrating the use of coupled Boltzmann equations for the Scalar Singlet model

### **Problem IV:**

Copy examples/aux/ScalarSinglet\_RD\_cBE.f to your private directory

- 1. Make sure that you can compile and run this main program
- 2. Try to understand what the program does, and how
- 3. Produce a plot corresponding to one of the curves in Fig. 3 of <u>1706.07433</u>.
- 4. Try to modify the program such that it can be linked to the generic\_wimp module instead
- 5. Replace generic\_wimp/kd/dskdm2.f and generic\_wimp/an/ dsanwax.f such that you get a difference between cBE and the standard approach
- **Hints :** The simplest way to ensure early kinetic decoupling is to make dskdm2 return just a power law in the SM particle energy, with an (ad hoc) small coefficient. As for dsanwx, you could e.g. multiply a Breit-Wigner resonance to the present output.